

- c. Suggest purchase of Westfall's *Construction of Modern Science* to provide you a reference book for background
- 3. Course requirements: a short, ungraded initial paper, three 6 to 8 page papers
 - a. Drafts of the three main papers on indicated dates, with final versions of the papers at the end of exam period
 - b. Papers play a critical role in learning in the course, and hence will insist within reason that drafts come in on time -- e.g. lose rewrite privilege otherwise, and receive little comments
 - c. Short initial paper to help you sort out the three different world systems
- 4. Also included with syllabus: a list of key events in science, the figures responsible, and their dates to give you a panoramic overview of the 3 centuries
 - a. At the end of each set of notes is a list of "select sources," the sources I have used in preparing the notes over the last 28 years; do not mistake this for an exhaustive bibliography
 - b. A valuable general reference: *Dictionary of Scientific Biography*
- 5. Three final comments about the course
 - a. Science distribution credit automatic for second semester, but not for first
 - b. A lecture course, but not in sense that you sit passively taking notes on what I say
 - (1) Lectures serve role of putting you in position to read and to write papers more effectively
 - (2) This by putting certain things hopefully in better focus and by providing you additional background
 - (3) Strongly encourage interruption and questions
 - (4) Notes will be available on Trunk from each class, plus optional further readings
 - c. A personal comment: there is no course I teach or have thought of teaching that I think has greater intellectual value for you
 - (1) Regardless of orientation -- science, engineering, humanities -- material of this course can provide a picture of science that will be of value the rest of your lives
 - (2) I intend the course to be one of the most memorable intellectual experiences of your life -- as it has proved to be for many people in it over the last 25 years

II. Astronomy, as Inherited by Ptolemy

A. The Celestial Sphere

- 1. The sphere of the stars: diurnal motion
 - a. Picture all the stars as if on a sphere surrounding the earth -- the celestial sphere (spherical astronomy)
 - b. Diurnal rotation about diameter through north and south celestial poles
 - c. Each star traces a circle, in part visible with time lapse photography
 - d. (The question of the distance of the stars, or any other celestial object, from the earth is the first fundamental evidence problem to note in this course!)

2. Celestial equator: the great circle midway between the two poles and perpendicular to the diameter through them
 - a. Instead of longitude and latitude, the corresponding "coordinate" system for the celestial sphere has right ascension and declination
 - b. Right ascension in units of hours, with 24 hours around the equator: corresponding to times when stars etc. ascend; declination in degrees, corresponding to geographical latitude
 3. In addition to the diurnal motion which it shares with the stars, the sun moves about 1 deg eastward each day with respect to the stars
 - a. Taking roughly 365 and 1/4 days to complete its way around the sphere
 - b. North of the celestial equator for roughly half of this time, south the other half
 4. The path the sun follows among the stars is called the ecliptic -- the line along which solar and lunar eclipses occur
 - a. Inclined at a little more than 23 deg from equator, with intersection points called the vernal and autumnal equinoxes, and the extreme points the two solstices
 - b. Celestial latitude and longitude measured with respect to the ecliptic, with a north and south ecliptic pole, and not with respect to the equator
 - c. From time immemorial, ecliptic divided into 12 constellations, each 30 deg wide: the signs of the zodiac
 5. The vernal equinox slowly moves among the stars, taking 26,000 years to complete its circuit
 - a. Precession of the equinoxes, actually from the wobble of the spheroidally shaped earth
 - b. From Aries to Pisces, soon to be entering the age of Aquarius
 - c. That precession occurs was known for a long time before the pattern and rate were determined; the precise cause was first announced in Newton's *Principia*
- B. Planets: "Wandering Stars"
1. The Sun is accordingly a wandering celestial body, in contrast to the "fixed stars", which stay in the same position with respect to one another
 - a. Wanders at a certain average rate per day
 - b. But in fact this rate not uniform: e.g. takes 94 and 1/2 days from the vernal equinox to the summer solstice, and 92 and 1/2 days from the summer solstice to the autumnal equinox
 2. The Moon is another example of a celestial body that "wanders" among the stars, at a much faster rate than the sun (around 27 and 1/3 days to complete the circuit), moving north and south of the ecliptic
 - a. Latitude of the moon quite complex: line of nodes itself moving with time
 - b. Eclipses as the classic prediction problem in astronomy, made difficult by the non-uniform longitudinal and latitudinal motion of the moon: the 18 year Saros cycle, a repeating cycle of sequences of eclipses, discovered by the Babylonians