

# NEWTON'S PROBLEM: TO INFER FORCES FROM MOTIONS

The *centripetal force* retaining a body in uniform circular motion varies as

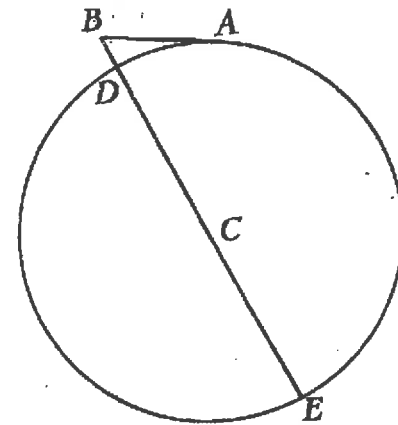
$$BD/\delta t^2$$

which – by Euclid 3,36 – becomes

$$\propto (AB^2/BE)/\delta t^2$$

which, as **B** approaches **A**,

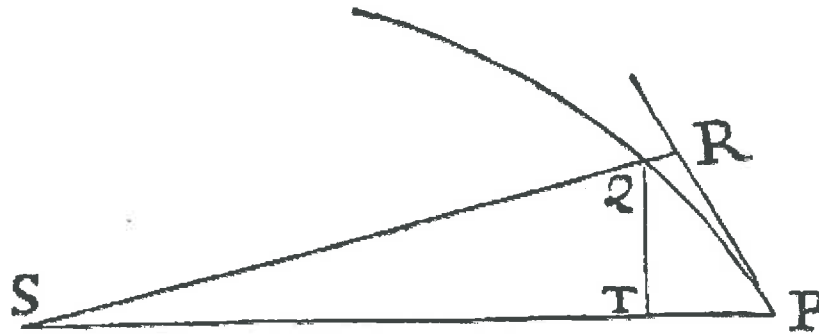
$$\propto v^2/r \propto r/P^2$$



**Problem:** How to generalize from uniform circular to arbitrary curvilinear motions – e.g. Kepler's ellipse?

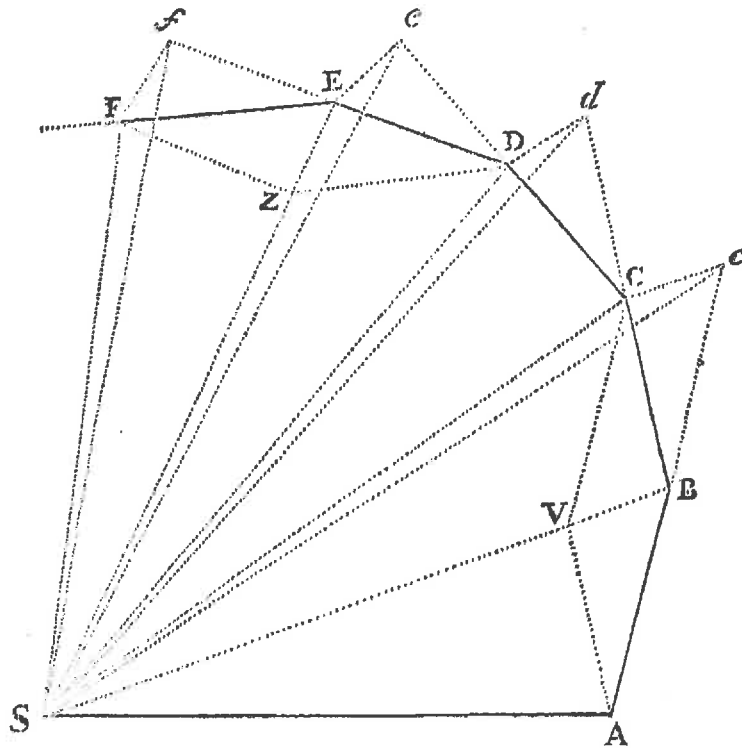
## Hooke's Proposal: Consider Only Forces Directed Toward a Single Point in Space

*Prop. 6:* To infer the “centripetal” force, toward S, from the curvilinear motion:



$$\text{force} \propto \lim \frac{QR}{\delta t^2}$$

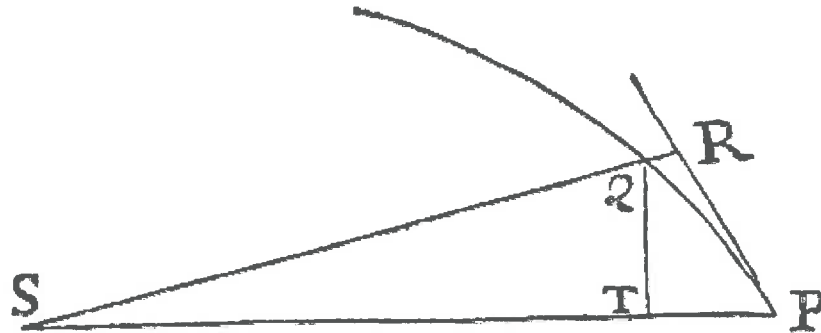
## The Area Rule: the First Key



**Theorem 1:** If all departures of a body from uniform motion in a straight line are directed toward a single point in space  $S$  – i.e. the external force on the body is centripetal – then the body sweeps out equal areas in equal times with respect to  $S$ .

# Newton's Basic Solution for Motion Under Centripetal Forces: the Second Key

*Prop. 6:* To infer the centripetal force, toward S, from the curvilinear motion:



$$\text{force} \propto \lim \frac{QR}{\delta t^2}$$

$$\propto \lim \frac{QR}{(QT^2 \times SP^2)}$$

$$\frac{1}{2r^2} \left( \frac{1}{r} + \frac{d^2}{d\theta^2} \left( \frac{1}{r} \right) \right)$$